

# MODELING OF 10 GEV-1 TEV LASER-PLASMA ACCELERATORS USING LORENTZ BOOSTED SIMULATIONS\*

J.-L. Vay<sup>†</sup>, C. G. R. Geddes, E. Esarey, W. Leemans, C. B. Schroeder, LBNL, USA  
D. P. Grote, LLNL, USA  
E. Cormier-Michel, Tech-X, USA

In a laser plasma accelerator, a laser pulse is propagated through a plasma, creating a wake of regions with very strong electric fields of alternating polarity [1]. An electron beam that is injected with the appropriate phase can thus be accelerated to high energy in a distance that is much shorter than with conventional acceleration techniques [2]. The simulation of a laser plasma acceleration stage from first principles using the Particle-In-Cell technique in the laboratory frame is very demanding computationally, as the evolution of micron-scale laser oscillations needs to be followed over millions of time steps as the laser pulse propagates through a meter-long plasma for a 10 GeV stage.

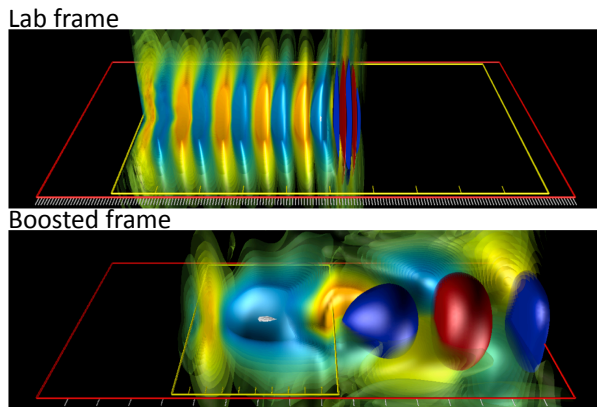


Figure 1: Simulations with the code Warp of scaled laser plasma acceleration stages: (top) in the lab; (bottom) in a Lorentz boosted frame (laser pulse in blue/red; plasma wakefield in pale blue/yellow).

A method was recently demonstrated to speed up full PIC simulations of a certain class of relativistic interactions by performing the calculation in a Lorentz boosted frame [3], taking advantage of the properties of space/time contraction and dilation of special relativity to render space and time scales (that are separated by orders of magnitude in the laboratory frame) commensurate in a Lorentz boosted frame, resulting in far fewer computer operations. As illustrated in Fig. 1, which shows snapshots from simulations of a sample LPA stage, in the laboratory frame the laser pulse is much shorter than the wake, whose wavelength is also much shorter than the acceleration distance ( $\lambda_{laser} \ll \lambda_{wake} \ll \lambda_{acceleration}$ ). In a Lorentz boosted frame co-propagating with the laser at a speed near the

speed of light, the laser is Lorentz expanded (by a factor  $(1 + v_f/c)\gamma_f$  where  $\gamma_f = (1 - v_f^2/c^2)^{-1/2}$  and  $v_f$  is the velocity of the frame and  $c$  is the speed of light). The plasma (now moving opposite to the incoming laser at velocity  $-v_f$ ) is Lorentz contracted (by a factor  $\gamma_f$ ). In a boosted frame moving with the wake ( $\gamma_f \approx \gamma_{wake}$ ), the laser wavelength, the wake and the acceleration length are now commensurate ( $\lambda_{laser} < \lambda_{wake} \approx \lambda_{acceleration}$ ), leading to far fewer time steps by a factor  $(1 + v_f/c)^2\gamma_f^2$ , hence computer operations [3, 4].

Recently, control of a violent numerical instability that limited early attempts [5, 6, 7] was obtained via the combination of: (i) the use of a tunable electromagnetic solver and an efficient wideband digital filtering method [8], (ii) observation of the benefits of hyperbolic rotation of space-time on the laser spectrum in boosted frame simulations [9], and (iii) identification of a special time step at which the growth rate of the instability is greatly reduced [8]. A novel numerical method for injecting the laser pulse through a moving planar antenna was also introduced [4]. The combination of these methods enabled the demonstration of a speedup of over a million times for the modeling of a hypothetical 1 TeV stage, and over 10,000 for a 10 GeV stage [9].

## REFERENCES

- [1] T. Tajima, J. Dawson, Phys. Rev. Lett. 43 (4) (1979) 267–270.
- [2] W. P. Leemans, B. Nagler, A. J. Gonsalves, C. Toth, K. Nakamura, C. G. R. Geddes, E. Esarey, C. B. Schroeder, S. M. Hooker, Nature Phys. 2 (10) (2006) 696–699.
- [3] J.-L. Vay, Phys. Rev. Lett. 98 (13) (2007) 130405/1–4.
- [4] J.-L. Vay, C. G. R. Geddes, E. Esarey, C. B. Schroeder, W. P. Leemans, E. Cormier-Michel, D. P. Grote, Phys. Plasmas 18 (2012) 123103.
- [5] D. Bruhwiler, J. Cary, B. Cowan, K. Paul, C. Geddes, P. Muldowney, P. Messmer, E. Esarey, E. Cormier-Michel, W. Leemans, J.-L. Vay, AIP Conference Proceedings, Vol. 1086, 2009, pp. 29–37.
- [6] J.-L. Vay, W. M. Fawley, C. G. R. Geddes, E. Cormier-Michel, Proc. Particle Accelerator Conference, Vancouver, Canada, 2009, TU1PBI04.
- [7] S. F. Martins, R. A. Fonseca, W. Lu, W. B. Mori, L. O. Silva, Nature Phys. 6 (4) (2010) 311–316.
- [8] J.-L. Vay, C. G. R. Geddes, E. Cormier-Michel, D. P. Grote, J. of Comput. Phys. 230 (15) (2011) 5908–5929.
- [9] J.-L. Vay, C. G. R. Geddes, E. Cormier-Michel, D. P. Grote, Phys. Plasmas 18 (3) (2011) 030701.

\* Supported by the US-DOE under Contract DE-AC02-05CH1123 and the SciDAC/ComPASS project. Used resources of the National Energy Research Supercomputer Center (NERSC).

<sup>†</sup> jlvay@lbl.gov