EMITTANCE MEASUREMENT ALGORITHM AND APPLICATION TO HIMM CYCLOTRON*

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Abstract

HIMM, a Heavy Ion Medical Machine, developed by Institute of Modern Physics, has been in operation since April 2020. The beam emittance of the cyclotron exit is measured with the most often used techniques, i.e. slit-grid, Q-scan and 3-grid at a dedicated beam line which is not the actual HIMM optical line. The high speed data acquisition architecture is based on FPGA, and motion control system is constructed based on the NI module.

The data post processing and emittance calculation is based on Python code with self-developed algorithm, including Levenberg–Marquardt optimization algorithm, thick lens model, dispersion effect correction, error bar fit, mismatch check, image denoise and "Zero-thresholding" calculation. The algorithm description and simulation are discussed in detail. The application of the algorithm to HIMM cyclotron is presented in this paper as well.

INTRODUCTION

The heavy ion medical machine (HIMM) is developed by the Institute of Modern Physics (IMP), which consists of two electron cyclotron resonance (ECR) ion sources, a cyclotron, a synchrotron ring and five nozzles [1]. The synchrotron has a compact structure with a circumference of 56.2 m. The layout of the HIMM complex is shown in Fig. 1. Up to now, the slow-extraction efficiency of HIMM has reached nearly 90% for all energies from 120 to 400 MeV/u. The spill duty factor has exceeded 90% at a sample rate of 10 kHz with the feedback-based slow-extraction technique applied [2, 3].



Figure 1: Layout of the HIMM complex.

Cyclotron, as the injector, plays a critical role in HIMM complex, which delivers the high quality beam to the ring through the MEBT (median energy beam transport line). The emittance and TWISS parameters measurement at the

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exit of the cyclotron are essential. Before integrating the cyclotron into HIMM, a dedicated beam line is constructed temporarily at the laboratory, and the measurement of emittance is performed along it.

The emittance and TWISS parameters at the exit of the cyclotron are measured using three most commonly used methods, slit–grid [4-6], Q-scan [7-9] and 3-grid [8-10]. A cross-check is of capability to validate this measurement.

In this paper, optimizition algorithm for Q-scan and 3grid is introduced, in which dispersion correction and thick lens model fit are of the most importance in improving the reconstruction accuracy of the emittance. In addition, slitgrid algorithm, especially denoising in reconstrucing the beam phase space is interpreted also. Both the optimizition algorithm and denoising process can result in a positive improvement for emittance calculation from the view point of simulation and analytic derivation. Finally, the measurement results of the cyclotron are presented.

ALGORITHM DESCRIPTION AND SIMULATION

Optimizition Algorithm for Q-scan and 3-grid

In the approximation of linear, beam transport along a lattice can be expressed in the form of matrix, which indicate how the transfer elements impose a function to the beam. From the view of matrix point, a transfer matrix can be formulated as

$$M = \begin{pmatrix} a & b \\ c & d \end{pmatrix} \tag{1}$$

without the loss of generality. The matrix must be symplectic, which is a remarkable property of a Hamiltonian system. After the implementation of the matrix, the squared distribution in the real space at the exit of the lattice element can be written as

$$\sum_{11}^{f} = a^{2} \left(\langle x_{\beta}^{2} \rangle + D^{2} \langle \delta^{2} \rangle \right) + 2ab \left(\langle x_{\beta} x_{\beta}' \rangle + DD' \langle \delta^{2} \rangle \right) + b^{2} \left(\langle x_{\beta}'^{2} \rangle + D'^{2} \langle \delta^{2} \rangle \right)$$
(2)

, with the dispersion considered and chromaticity ignored. For a typical configuration of focus lens followed by a drift, the a and b can be expressed as

$$a(k) = \cos\sqrt{kl} - \sqrt{kL}\sin\sqrt{kl}$$

$$b(k) = \frac{1}{\sqrt{k}}\sin\sqrt{kl} + L\cos\sqrt{kl}$$
 (3)

with *L* the drift length, $k = B'/B\rho$, $B'l = C_1 + C_2 I$, C_1 and C_2 for calibrated coefficient of magnet. For Q-scan scheme, the desired values, i.e., $\langle x_{\beta}^2 \rangle$, $\langle x_{\beta} x_{\beta}' \rangle$, $\langle x_{\beta}'^2 \rangle$, can be obtained via solving a non-linear least squares problem with the objective function given as follows

$$\sum_{11}^{f} (\boldsymbol{I}, \boldsymbol{\alpha}; \boldsymbol{D}, \boldsymbol{D}', \boldsymbol{\delta}) = \sum_{11}^{f}$$
(4)

, in which, $\boldsymbol{\alpha} = (\langle x_{\beta}^2 \rangle, \langle x_{\beta} x_{\beta}' \rangle, \langle x_{\beta}'^2 \rangle)$. To find a solution, the following minimization procedure would be implemented,

$$= \arg \min_{\alpha} S(\alpha) \tag{5}$$

where
$$S(\boldsymbol{\alpha}) = \sum_{i=1}^{n} \left[y_i - \sqrt{\sum_{1=1}^{f} (\boldsymbol{I}, \boldsymbol{\alpha}; \boldsymbol{D}, \boldsymbol{D}', \delta)} \right]^2$$
.

α

If data uncertainties involved, the minimization may be scaled with a weighting factor ε_i that refers to the uncertainties of the profile y_i ,

$$\boldsymbol{\alpha} = \arg\min_{\boldsymbol{\alpha}} \sum_{i=1}^{n} \left[\frac{y_i - \sqrt{\sum_{11}^{f} (\boldsymbol{I}, \boldsymbol{\alpha}; \boldsymbol{D}, \boldsymbol{D}', \boldsymbol{\delta})}}{\varepsilon_i} \right]^2 \qquad (6)$$

The most widely used optimization algorithm for this is the so called Levenberg–Marquardt algorithm (LMA or just LM), which can be thought of a combination of the method of gradient descent and the Gauss-Newton algorithm (GNA), both of which are the most classical optimization algorithm. Thus, LMA adapts the advantages of both of the two algorithm, and hence has high performance as expected [11-13].



Figure 2: The difference between parabolic fit based thin lens model and LMA fit based thick lens model. Red dotline is parabolic fit based thin lens model. Black dot-line is LMA fit based thick lens model, which agrees well with the theoretical emittance depicted by blue dot-line.

If (D, D', δ) are given at the measurement point beforehand, the TWISS parameters can be extracted with a high accuracy. For the case of unknown (D, D', δ) , to extract the whole parameters including (D, D', δ) simultaneously, at least one dipole must be involved along the transport line.

For 3-grid scheme, the variable would be drift length L instead, hence, the objective function will be

$$\sum_{11}^{f} (\boldsymbol{L}, \boldsymbol{\alpha}; \boldsymbol{D}, \boldsymbol{D}', \boldsymbol{\delta})$$
(7)

LMA is an iterative approach with trust-region searching strategy. For each iteration, the Jacobian matrix $J = \partial r / \partial \alpha$ is given as the first step, and then the advance of α , $\Delta \alpha$, is found by solving the following linear equation

$$(\boldsymbol{J}^T\boldsymbol{J} + \lambda\boldsymbol{I})\Delta\boldsymbol{\alpha} = -\boldsymbol{J}^T\boldsymbol{r}$$
(8)

Where I is the identity matrix and λ is a damping parameter to control the solution $\Delta \alpha$ brings $r^T r$ down hill how fast. The error of the fitting parameters is estimated by the covariance matrix defined as $C = J^T J$ and then the 1σ error bar is $\sigma_i = \sqrt{C_{ii}}$ for the i'th paremeter [14].

A simulaton is performed based on the discussion above. First, the dispersion-free beam line is tracked, comparision between parabolic fit based thin lens model and LMA fit based thick lens model is presented in Fig. 2. With thin lens model and parabolic fit applied, the solutions don't converge to the actual one, and a visible difference is observed as expected. Highly overestimated results occur at high current cases, which can be inferred from the fact that thin lens approximation is reasonable only when the quadrupole current is low. LMA fit based thick lens model gives the desired results in comparison to the actual one with a deviation of 1.56%.

Typically, dispersion can result in overestimated measurement of emittance from the enlarged beam width. The typical treatment for dispersion is to subtract its contribution from the measured beam width [15], like that $\langle x_{\beta}^2 \rangle = \sqrt{\langle x_{meas}^2 \rangle - D^2 \langle \delta^2 \rangle}$. However, this procedure is sometimes error-prone and of less robustness since the dispersion has to be evaluated explicitly as a function of quadrupole settings. An alternative is to take account of it as the fixed parameters in the fit model, which is more reasonable if the dispersion is given at the initial position as the first time, as interpreted in Eq. (5).



Figure 3: Left: non-zero dispersion induced beam width change. (green star) Diagnostics section is dispersion-free. (blue angle) Residual dispersion D = 5 m at the entrance of quadrupole. (red line) Fit using thick lens model with dispersion correction algorithm applied. Right: reconstructed phase space plotted in normalized coordinates frame, which displays the mismatch qualitatively.

The simulation results of emittance change induced by dispersion are shown in Figs. 3 and 4. The beam width change induced by a non-zero dispersion can be observed

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from Fig. 3 left. With the increase of dispersion, the reconstructed emittance is overestimated more and more, especially when the derivative of dispersion is positive, as presented in Fig. 4. If fit model treats dispersion as known parameters as formulated in Eq. (5), the dispersion-free TWISS parameters can be achieved with a good accuracy. This can be observed alternatively from the normalized phase space illustrated in the right plot of Fig. 3. which displays the mismatch between the designed (machine circle) and measured phase space (measured circle). A mismatch factor can be calculated as described in [16, 17],

$$\xi = \frac{1}{2} (\gamma_D \beta - 2\alpha_D \alpha + \beta_D \gamma) \tag{9}$$

if the designed TWISS (β_D, γ_D) is equal to the measured one (β, γ) , the beam is matched to the machine optical, otherwise the mismatch is indicated by ξ .



Figure 4: The comparison between with and without dispersion correction applied. Before correction and dispersion derivative D' > 0 (red), before correction and dispersion derivative D' < 0 (black), after correction (blue), theoretical one (purple).

'Zero-Thresholding' Approach for Slit-Grid Technique

The slit-grid measurement would generate a 2D phase space that is the area the emittance calculation launched. From which, the RMS emittance can be retrieved statistically. To simplify the calculation, the data index of the phase space matrix are used instead of the data itself [18]. The RMS emittance is defined as

$$\varepsilon_{rms} = \sigma_{n_x} \sigma_{n_{px}} \Delta x \Delta x' \sqrt{1 - \rho^2}$$
(10)

the notation is the same with the reference [18].

Typically, the data are noisy owing to disturbances from the MW electronics and the environment. It is, therefore, necessary to denoise the data before performing the calculation. A systematic procedure to complete the emittance evaluation is given in [19], in which, a method called 'zero-thresholding' approach is proposed. For the method, the beam phase space is processed following a systematic procedure, including ad hoc-based interpolation, wavelet-based denoising, threshold-based ROI selection, and "zero-thresholding"-based emittance extraction. From the simulation, the quantized error decreases with the increased interpolation number, and a deviation of below 4% is reached when interpolation number 4 is given. The theory supporting the "zero-thresholding" based emittance extraction is given in APPENDANCE.

APPLYING ALGORITHM TO HIMM CYCLOTRON

The emittance of the HIMM cyclotron is measured and the data processing is finished. Three techniques are employed to complete this procedure, and the algorithm disscused above is applyed. The results are shown in Table 1, Figs. 5-7.

The results of these measurements agree well with each other for the vertical plane. There is a discrepancy for the horizontal plane, however. Based on our deductions, this is attributed to the nonzero dispersion contribution. The dispersion and its derivative are not known with a high accuracy in the horizontal plane at the exit of cyclotron, and the temporal beam line is not designed to be dispersion-free in the diagnostics section.

Table 1: Emittance Summary at the Exit of HIMM Cyclotron Measured Using Three Techniques

	Slit-grid	Q_scan	3-grid
Horizontal(95%) [mm*mrad]	45.32	51.5	56.93
Vertical (95%) [mm*mrad]	37.56	31.8	38.15

APPENDANCE

For simplicity, assumping the real space distributuion obeys 1D Gaussian function with zero mean, that is $x \sim \mathcal{N}(x; 0, \sigma)$. The probability density function can be formulated as

$$f(x;\sigma) = \frac{1}{\sqrt{2\pi\sigma}} \exp\left(-\frac{x^2}{2\sigma^2}\right).$$
(A1)

The maximum intensity locates at x = 0, and the value is $I_{max} = f(0) = 1/\sqrt{2\pi\sigma}$. Assumping the threshold is $I_{th} = \chi I_{max}$, in which, $\chi \in (0,1)$ means the proportion of the signal the threshold occupied. A larger χ is corresponding to a smaller phase space, and hence a cleaner phase space. The phase space coordinates x_{th} corresponding to the threshold I_{th} can be evaluated by $f(x_{th}) = I_{th}$, which gives

$$x_{th} = \pm \sqrt{-2ln(\chi\sigma)}.$$
 (A2)

If beam phase space is not cut by the threshold, that is $\chi = 0$ or $x_{th} = \pm \infty$, the statistic deviation of the distribution is σ exactly. However, when it is cut, the new statistic deviation should be evaluated via the definition as

$$\sigma_{th}^{2} = \langle (x - \langle x \rangle)^{2} \rangle = \frac{\int_{-x_{th}}^{x_{th}} x^{2} f(x) dx}{\int_{-x_{th}}^{x_{th}} f(x) dx} = \frac{-2\chi\sigma^{2}[iDaw(i\sqrt{-ln\chi}) + \sqrt{-ln\chi}]}{\sqrt{\pi}Erf(\sqrt{-ln\chi})}$$
(A3)

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Figure 5: Phase space distribution obtained from slit-grid technique. Left: the horizontal plane, Right: the vertical plane. Top: the raw phase space, bottom: the denoised phase space, the red dashed line is the contour line with a contour level of 10%, and the black dashed line is the measured 1σ phase ellipse.









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, in which *i* is the complex, $Daw(x) \equiv \exp(-x^2) \int_0^x \exp(t^2) dt$ is the Dawson's integral, $Erf(z) \equiv \frac{2}{\sqrt{\pi}} \int_0^z \exp(-t^2) dt$ is the Error function. The ratio between σ_{th} and σ is

$$\delta^{2}(\chi) = \frac{\sigma_{th}^{2}}{\sigma^{2}} = \frac{-2\chi[iDaw(i\sqrt{-ln\chi}) + \sqrt{-ln\chi}]}{\sqrt{\pi}Erf(\sqrt{-ln\chi})}.$$
 (A4)

The evolution of δ^2 with χ is shown in Fig. 8. It is a monotonic decreasing function. The limit of $\delta^2(\chi)$ near the $\chi = 0$ is

$$\lim_{\chi \to 0} \delta^2(\chi) = 1 \tag{A5}$$

, which indicates that σ_{th} approaches σ when χ approaches 0. It geives the correctness of the formulation Eq. (A4). With the increase of χ , a linear section occurs, which gives the possibility to approximate the σ with σ_{th} linearly. Series expansion of $\delta^2(\chi)$ at far away from $\chi = 0$ is

$$\begin{split} \delta^2(\chi) &= 0.823 - 1.453(\chi - 0.1) \\ &+ \mathcal{O}((\chi - 0.1)^2). \end{split} \tag{A6}$$

Linear approximation gives a good estimation of σ , as depicted in Fig. 8. Therein the dotted red line is a linear approximation of $\delta^2(\chi)$. The deviation is large somewhat near the $\chi = 0$. The maximum deviation is 0.03, and the maximum deviation in terms of statistic deviation σ is less than 2%.



Figure 8: The evolution of δ^2 with χ . Green line is $\delta^2(\chi)$, dotted red line is a linear approximation.

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