



# EIC Beam Dynamics Challenges

Derong Xu on behalf of EIC Design Team

Brookhaven National Laboratory  
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Electron-Ion Collider

# Outline

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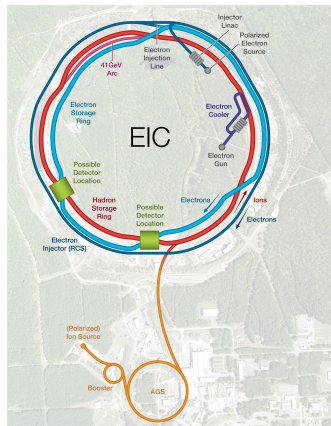
# Introduction — Electron Ion Collider

## Science goals

- How does the mass of the nucleon arise?
- How does the spin of the nucleon arise?
- What are the emergent properties of dense systems of gluons?

## Design goals

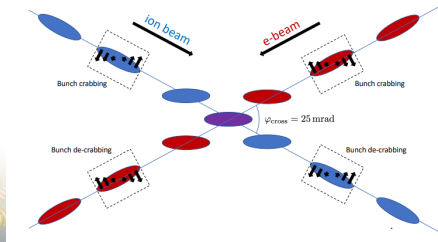
- High luminosity:  $10^{33} - 10^{34} \text{ cm}^{-2}\text{s}^{-1}$
- center-of-mass energies: 20 – 140 GeV
- Polarized proton and electron beams: 70%
- Large range of hadron species: Proton - Uranium
- Possibility of 2nd IR



HSR — Hadron Storage Ring  
ESR — Electron Storage Ring  
RCS — Rapid Cycling Synchrotron

# Luminosity — Overview

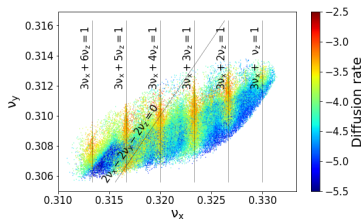
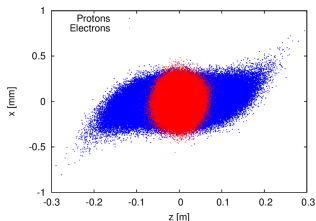
- Large crossing angle 25 mrad, fast separation to avoid parasitic collision
- Local crab crossing: upstream and downstream crab cavities to restore effective head-on collision to compensate geometric luminosity loss
- Large beam-beam parameters,  $e \sim 0.1, p \sim 0.015$ , combination never experimentally demonstrated
- Flat beam  $\sigma_y/\sigma_x = 0.09$  to achieve highest e-p luminosity  $10^{34} \text{ cm}^{-2}\text{s}^{-1}$



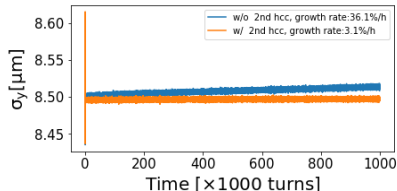
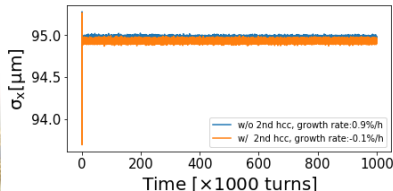
Parameter	unit	proton	electron
Circumference	m	3833.8451	
Particle energy	GeV	275	10
Bunch intensity	$10^{11}$	0.668	1.72
# of Bunches	-	1160	
Crossing angle	mrad	25	
$\beta^*$ at IP	cm	80/7.2	45/5.6
Beam sizes at IP	$\mu\text{m}$	95/8.5	
Bunch length	cm	6	2
Energy spread	$10^{-4}$	6.6	5.5
Transverse tunes	-	0.228/0.210	0.08/0.06
Longitudinal tune	-	0.01	0.069
BB parameter	-	0.012/0.012	0.07/0.10
Luminosity	$\text{cm}^{-2}\text{s}^{-1}$		$10^{34}$

# Luminosity — Synchro-betatron resonances (SBR)

- Crabbed offset, due to sinusoidal kick from the crab cavities, drives higher-order SBR in the proton beam  $\Delta x = -\theta_c [\sin(k_c z)/k_c - z]$
- $\nu_x = 0.228, \nu_y = 0.210$  to mitigate SBR

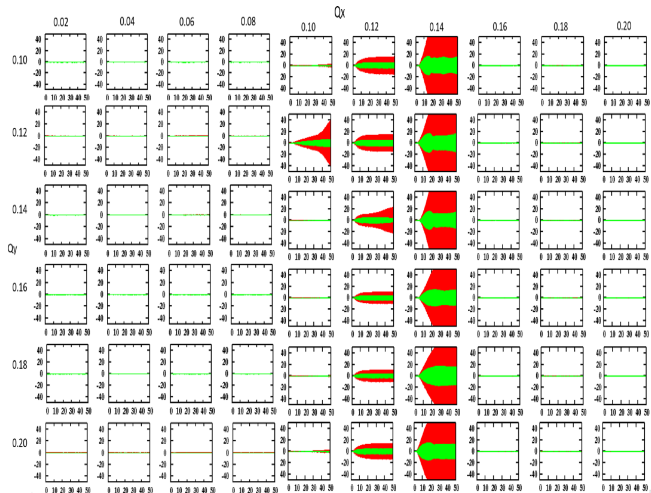


- 2nd order harmonic crab cavity is used to flatten the crabbed offset



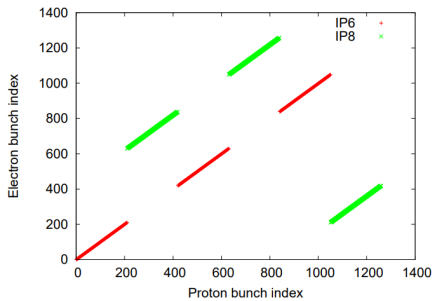
# Luminosity — Coherent beam-beam effects

Electron tune scan to avoid coherent instability, red: proton H centroid, green: electron H centroid. Electron working point: (0.08, 0.06)



# Luminosity — Collision with two IPs

- A 2nd IR is reserved at IP8 for future upgrade. However, with same beam-beam parameters, sum luminosity  $\propto 1/N_{IP}$
- Time sharing, one IR taking data while the other one is idle
- Or **luminosity sharing**, change bunch filling pattern, so that half bunches collide at IP6, and the other bunches collide at IP8

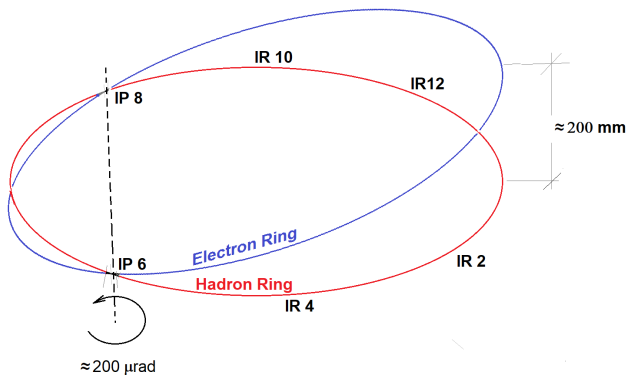


- The other half electron bunches are shifted by 3 RF buckets
- IP8 should be moved away from IP6 for synchronization
- Each bunch only collides **once per turn**

# Luminosity — Tilted ESR

HSR lies in horizontal plane, while ESR is tilted by  $\sim 200 \mu\text{rad}$

- Resolve interferences between rings, transfer lines, cooler ERL in IR2
- Avoid vertical bends around ring crossing points to preserve polarization
- Effect on dynamics:  $-4 \text{ mrad}$  rotation around  $s$  axis before collision, and  $4 \text{ mrad}$  rotation after collision — vertical crabbing needed





# Polarization — Overview

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The EIC physics program requires highly polarized hadron and electron beams with alternating spin orientation for the electron bunches

Polarized hadron beam:

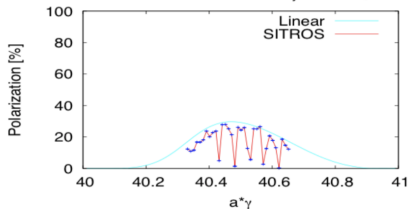
- Improvements of AGS: 70% achieved at extraction
- Four additional Siberian snakes will be installed in HSR:
  - increase polarization transmission for protons on the ramp to 275 GeV to  $\sim 100\%$
  - sufficiently suppress spin resonance width for the polarized  $^3\text{He}$
- spin rotators based on helical dipoles to transform spin directions

Polarized electron beam:

- 85% longitudinal polarization in the source
- RCS serves as acceleration and injection at full energy, high periodicity to be free of intrinsic spin resonances
- Frequent "**swap-out**" injection to keep time-averaged polarization

# Polarization — Depolarization in ESR

- Spin matching of spin rotator optics minimized beam depolarization, especially at 18 GeV
- Spin simulation studies with magnet errors showed that with one interaction region the average polarization of at least 70% is achievable. Studies with two IRs are underway
- Vertical emittance can be achieved using vertical bumps in ESR arcs, with depolarization at an acceptable level



Orbit corrected to rms:  $\sim 0.15$  mm, coupling corrected to below 0.005. SITROS includes nonlinear sextupole fields and quantum excitation

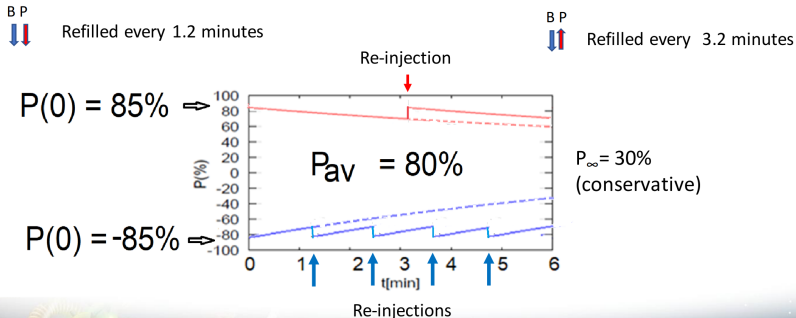
Assumed quadrupole RMS misalignments

horizontal offset	$\delta x^Q$	200 $\mu\text{m}$
vertical offset	$\delta y^Q$	200 $\mu\text{m}$
roll angle	$\delta\psi^Q$	200 $\mu\text{rad}$

At 18 GeV with 2.5 min refill time: 16% asymptotic polarization corresponds to 70% average polarization

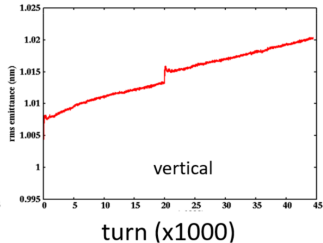
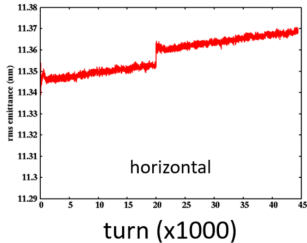
# Polarization — Electron bunch replacement

- Physics program requires bunches with spin “**up**” and spin “**down**” (in the arcs) to be stored **simultaneously**
- Initial polarization  $\sim 85\%$  decays towards  $P_\infty < \sim 50\%$ : Sokolov-Ternov self-polarization and spin diffusion
- Frequent injection is necessary to keep time-averaged polarization
- At 18 GeV, every bunch is replaced (on average) after 2.2 min with RCS cycling rate of 2Hz

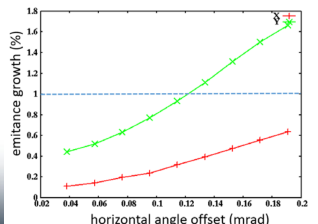
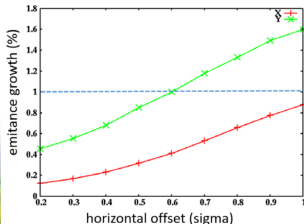


# Polarization — Electron bunch replacement

0.1% emittance growth when injection on orbit

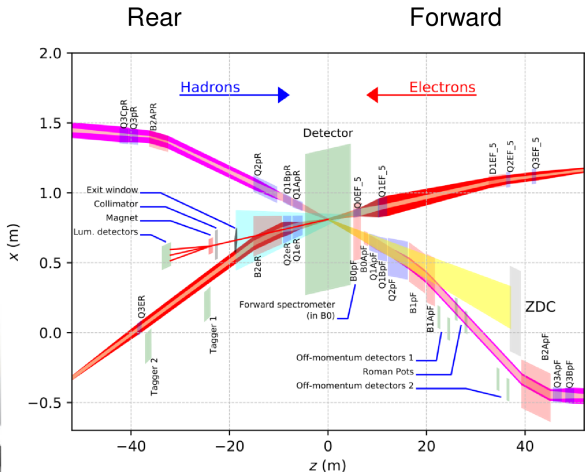


1% emittance growth corresponding to  $60\mu\text{m}$  or  $0.12\text{mrad}$  injection errors



# Linear Beam Optics — IR design

- Strong focusing at IP, HSR:80/7.2 cm, ESR:45/5.6 cm
- Crab cavities: high  $\beta_x$ , specific  $\Psi_x$ , and enough installation space
- Accommodation to detector: 4.5 m rear, 5.0 m forward stay-clear...



# Linear Beam Optics — Crab dispersion control

Crab cavities introduce  $z$ -dependent transverse kick



$$\zeta = \left( \frac{\partial x}{\partial z}, \frac{\partial x'}{\partial z}, \frac{\partial y}{\partial z}, \frac{\partial y'}{\partial z} \right)^T, \quad \zeta_2 = M\zeta_1$$

Crab dispersion closure

- Ideally, two thin crab cavities apart with  $n\pi$  phase advance form a **closed crab dispersion bump**
- In both rings, the crab cavities **can not** be matched to exactly  $\pi/2$
- In ESR, the bump has to be closed because  $\nu_x \approx \nu_z$ . This is accomplished by moving rear side crab to  $\sim 3\pi/2$  ( $2\pi$  between both crabs)
- In HSR, the crab dispersion bump is **not** closed ( $5^\circ$  away from  $\pi$ ). Crab cavity voltages can be adjusted to provide  $\zeta^* = (12.5 \text{ mrad}, 0, 0, 0)$
- Exploring the necessity of and options for locally closing crab dispersion
- Reasonable momentum dispersion constraints at crab cavities to reduce their combined effects

# Linear Beam Optics — Crab dispersion control

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## Vertical crabbing

- Sources: tilted ESR, detector solenoid
- Vertical crab cavities can provide knobs to control vertical crab dispersion. However, (1) hard to match, (2) conflict with impedance budget
- Skew quadrupoles are feasible and efficient to control vertical crabbing
- In ESR, the required skew component strength is 1.2 T/m. In HSR, it may combine with the global decoupling system.

## Dynamical control — crab cavity RF noise

- The crab dispersion changes dynamically due to RF phase and amplitude noise. The transverse emittance growth  $\propto \theta_c^2$
- Compared with HL-LHC, EIC sensitivity to RF noise is 4000 times higher; the emittance growth tolerance is 3 orders of magnitude higher
- The RF noise threshold for the HSR will be very **hard** to achieve. A dedicated feedback system is **needed**

# Linear Beam Optics — ESR spin rotation

The aim of the spin rotators is: (1) to rotate the spin from the vertical direction in the arcs to the longitudinal direction at the IPs; (2) to minimize beam depolarization.

ESR spin rotator composes of solenoids and dipoles



- In dipoles, spin rotated around vertical axis by  $\psi = a\gamma\theta$ . In solenoids, spin rotated around longitudinal axis by  $\varphi = (1 + a)KL$
- One "long" solenoid module for 18 GeV,  $\varphi_1 = 0, \psi_2 = \varphi_2 = \pi/2$
- One "short" solenoid module for 6 GeV,  $\varphi_1 = \psi_1 + \psi_2 = \pi/2$
- Both solenoid modules used for 10 GeV
- Matching for 5 GeV electrons is ongoing

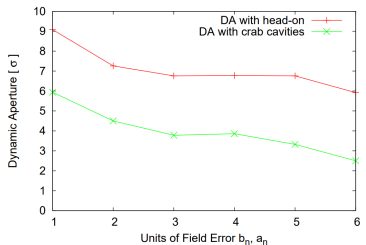
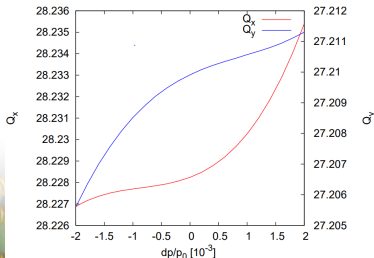


# Dynamic Aperture — HSR overview

- HSR will reuse arcs of both Yellow and Blue RHIC rings
- Sufficient DA after the linear chromaticity is corrected. More sextupole families are available for further DA optimization
- IR magnetic field errors dominate hadron ring DA reduction

$$\Delta B_y + i\Delta B_x = B(R_{\text{ref}}) \left[ 10^{-4} \sum_{n=0}^{N_{\text{max}}} (b_n + ia_n) \frac{(x + iy)^n}{R_{\text{ref}}^n} \right]$$

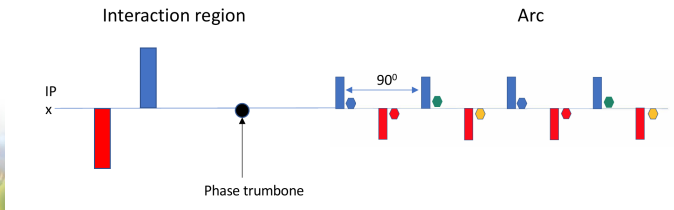
- Tracking with beam-beam:  $3\sigma$  drop from head-on to crab collision with 1 unit IR field errors (due to crab dispersion and IR field errors)



# Dynamic Aperture — ESR overview

Dynamic aperture and momentum aperture of lower energies lattice are sufficient. 18 GeV lattice with a 2nd IR is **most** challenging

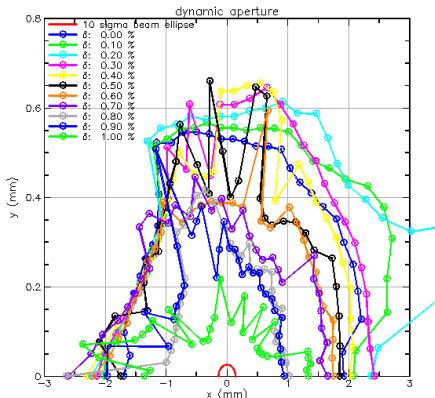
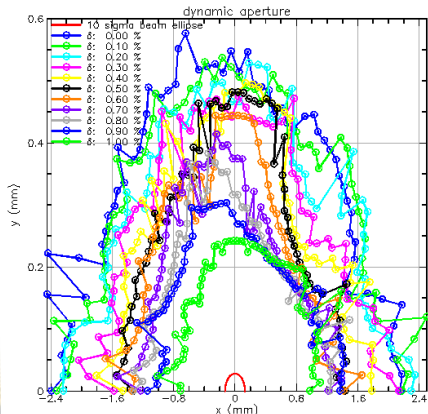
- Optimization goal:  $10\sigma$  in all three planes
- Fractional tunes close to integer: selected by spin and beam-beam dynamics
- Setting the phase advance between IRs to  $(2n + 1)\pi/2$  helps, but not sufficient
- No space for local chromatic compensation
- The off-momentum  $\beta$ -beating and the chromaticity from the final focusing doublet are corrected in the neighboring arc section



# Dynamic Aperture — ESR 18 GeV tracking results

Bare lattice

With beam-beam, crab cavities, detector solenoid, and crab dispersion correction by skew quadrupoles



10 $\sigma$  with  $\sigma_\delta = 0.1\%$  is **achieved** for the 18 GeV lattice with 2 IRs

# Collective Effects — Overview

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In ESR, average current: 2.5 A with bunch charge of 28 nC

- No single bunch instabilities, component heating needs water cooling
- With a 591 MHz RF system  $\sigma_z = 7$  mm,  $I_{\text{peak}} = 480$  A
- Large tune spread caused by beam-beam interaction provide Landau damping for transverse coupled-bunch instability and ion instability
- Longitudinal damper is needed

In HSR, average current: 1 A

- RHIC vacuum chamber is not designed for EIC beam
- Vulnerable to electron cloud instability and high resistive losses from beam-induced currents
- The vacuum chamber of the HSR SC magnets and their cold interconnects will be updated with a beam screen to present sufficiently **low** impedance and **low** secondary electron-emission yield (SEY)

# Collective Effects — ESR impedance budget

- We have initial designs for main components and wakefield calculations
- Beam is stable at single bunch current of 2.2 mA required for regular operation at 2.5 A within 1160 bunches

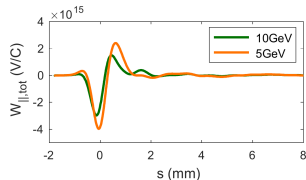
Components	Abbreviation	Number	Status
Bellows	BLW	350	✓ x2 (NEG)
Collimator Ramp <sup>1</sup>	CLM	16	✓
Horizontal In-Vacuum Collimator	HIVC	3	TBD
Vertical In-Vacuum Collimator	VIVC	3	TBD
Crab Cavity	CRBCVT	2	✓
Beam Position Monitor <sup>2</sup>	BPM	494	✓
Gate Valve <sup>2</sup>	GV	30	✓
Stripline Kicker <sup>2</sup>	SK	18	✓
Main RF Cavity <sup>2</sup>	CVT	23	✓
Tapered Transition in RF Section	TPRD	9?	TBD
Multipole Chamber Absorber	MPABS	292	✓
Dipole Chamber Absorber	DPABS	250	✓
Flange Joints	FLNG	1500	TBD
Resistive Wall	RW	-	✓

1 - SKEKB design

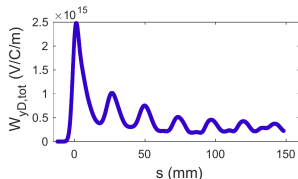
✓ - Included into the total W(s)

2 - NSLS-II design

The total longitudinal wakefield simulated for a 0.3 mm bunch length at 5 GeV and 10/18 GeV energies.

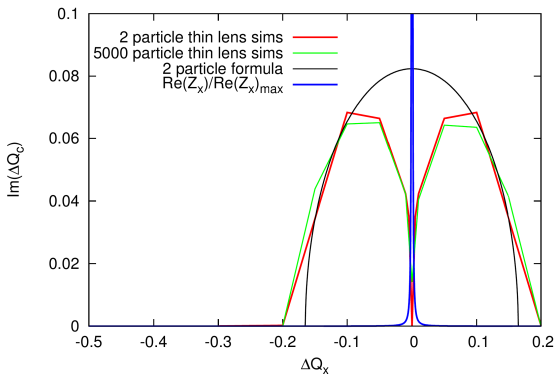


The total vertical dipole wakefield simulated for a 2 mm bunch length.



## Collective Effects — Fundamental crabbing mode

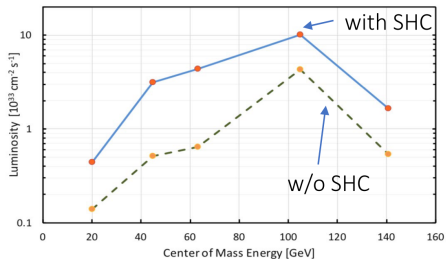
Transverse Crab cavities with big  $R/Q$  can lead to transverse coupled bunch instabilities: high  $Q \sim 10^6$ , and high  $\beta_x \sim 1300$  m in 275 GeV HSR



RF feedback is **required** on the crab cavities,  $Q_{\text{eff}} = 300$  for 197 MHz and  $Q_{\text{eff}} = 600$  for 394 MHz

# Strong Hadron Cooling — Overview

- Luminosity benefits strongly (factor  $\approx 3 - 10$ ) from cooling the transverse and longitudinal hadron beam emittance
- IBS longitudinal and transverse growth time is 2-3 hours. The cooling time shall be equal to or less than the growth time from all sources
- Cooling at 275 GeV and 100 GeV based on Coherent electron Cooling (CeC); 41 GeV cooling under study.
- Low energy cooling (Pre-cooling based on LEReC) is used to obtain initial parameters of proton beam: must cool the hadron beam normalized vertical emittance from  $2.5 \mu\text{m}$  to  $0.3 \mu\text{m}$  in 2 hours



## Summary — Challenges resolved and in progress

Challenges from design parameters: luminosity  $10^{34} \text{cm}^{-2} \text{s}^{-1}$  (factor 100 beyond HERA), both beams of time average polarization  $> 70\%$

- Beam-beam parameters, achievable in simulation
- Electron bunch replacement, sufficient injection errors
- Spin rotators design, 6 – 18 GeV, 5 GeV
- Strong hadron cooling

Challenges from crab cavities

- Higher-order SBR, tune optimization and 2nd-order harmonic
- Crab dispersion control, ESR, HSR closure
- RF feedback to cure RF noise and large transverse impedance
- HSR DA reduction, 1IR with reasonable IR field error

Challenges from 2 IRs

- ESR Chromatic correction,  $10\sigma$  achieved with 2 IRs
- Luminosity sharing, bunch filling pattern
- HSR DA, ESR depolarization...



## Acknowledgement

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*E. C. Aschenauer, G. Bassi, J. Beebe-Wang, S. Benson, J. S. Berg, W. Bergan, M. Blaskiewicz, A. Blednykh, J. M. Brennan, S. Brooks, K. A. Brown, Y. Cai, Z. Conway, J. R. Delayen, K. A. Drees, A. V. Fedotov, W. Fischer, C. Folz, B.R.P. Gamage, D. Gassner, E. Gianfelice-Wendt, J. Grames, X. Gu, R. Gupta, Y. Hao, C. Hetzel, G. Hoffstaetter, D. Holmes, H. Huang, H. Lovelace III, J. Kewisch, Y. Li, F. Lin, C. Liu, Y. Luo, G. Mahler, D. Marx, F. Meot, T. Michalski, M. Minty, C. Montag, V. Morozov, S. Nayak, E. Nissen, Y. Nosochkov, R. B. Palmer, B. Parker, S. Peggs, B. Podobedov, J. Preble, V. Ptitsyn, J. Qiang, V. H. Ranjbar, R. Rimmer, G. Robert-Demolaize, D. Sagan, M. Sangroula, T. Satogata, S. Seletskiy, A. Seryi, M. Signorelli, S. D. Silva, K. S. Smith, G. Stupakov, M. Sullivan, S. Tepikian, R. Than, P. Thieberger, N. Tsoupas, J. Tuozzolo, J. Unger, S. Verdu-Andres, E. Wang, D. Weiss, F. J. Willeke, M. Wiseman, H. Witte, W. Wittmer, Q. Wu, W. Xu, A. Zaltsman*

Thank you for your attention.